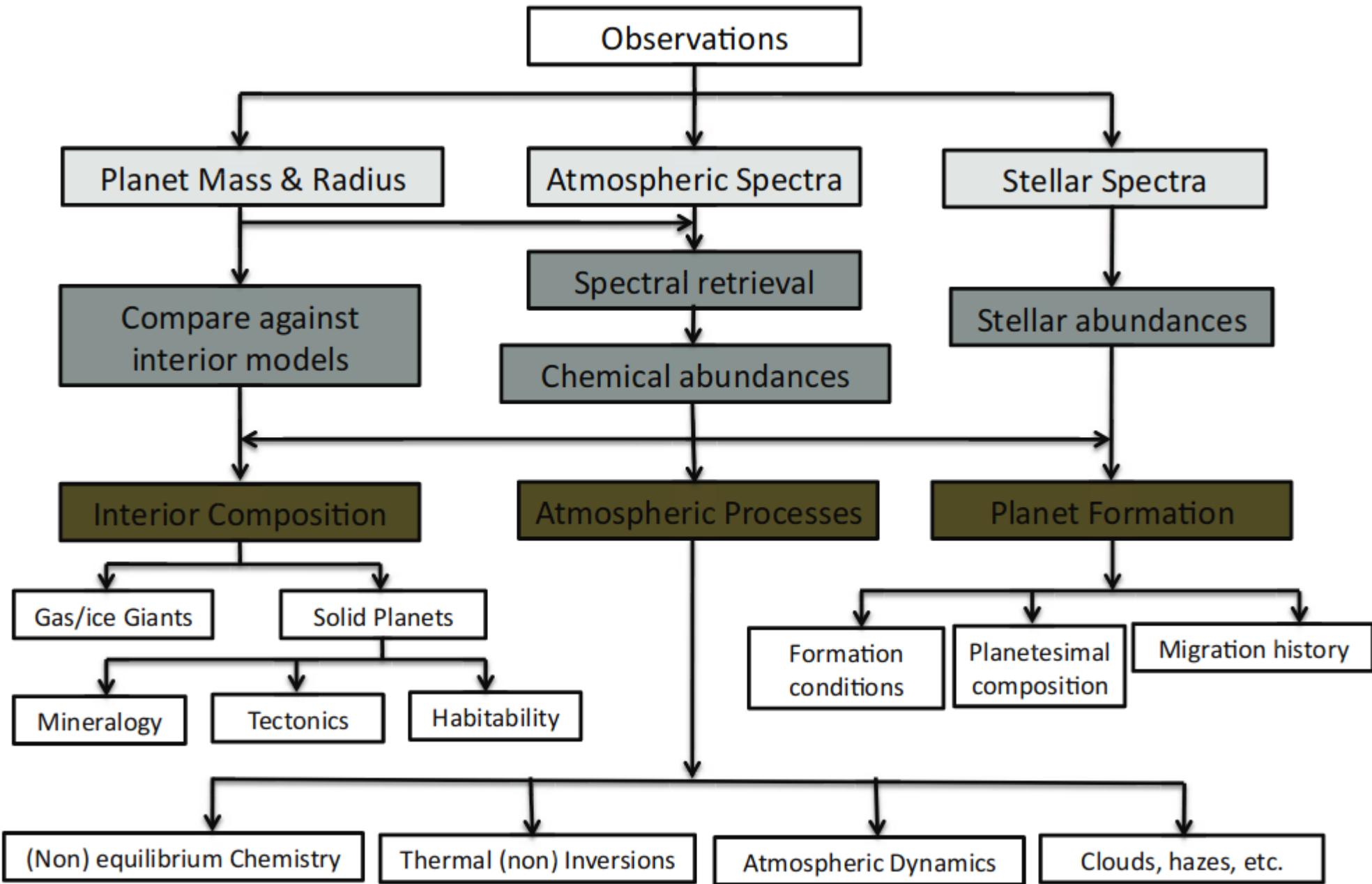


Exoplanet characterization

Exoplanet characterization



Exoplanet characterization

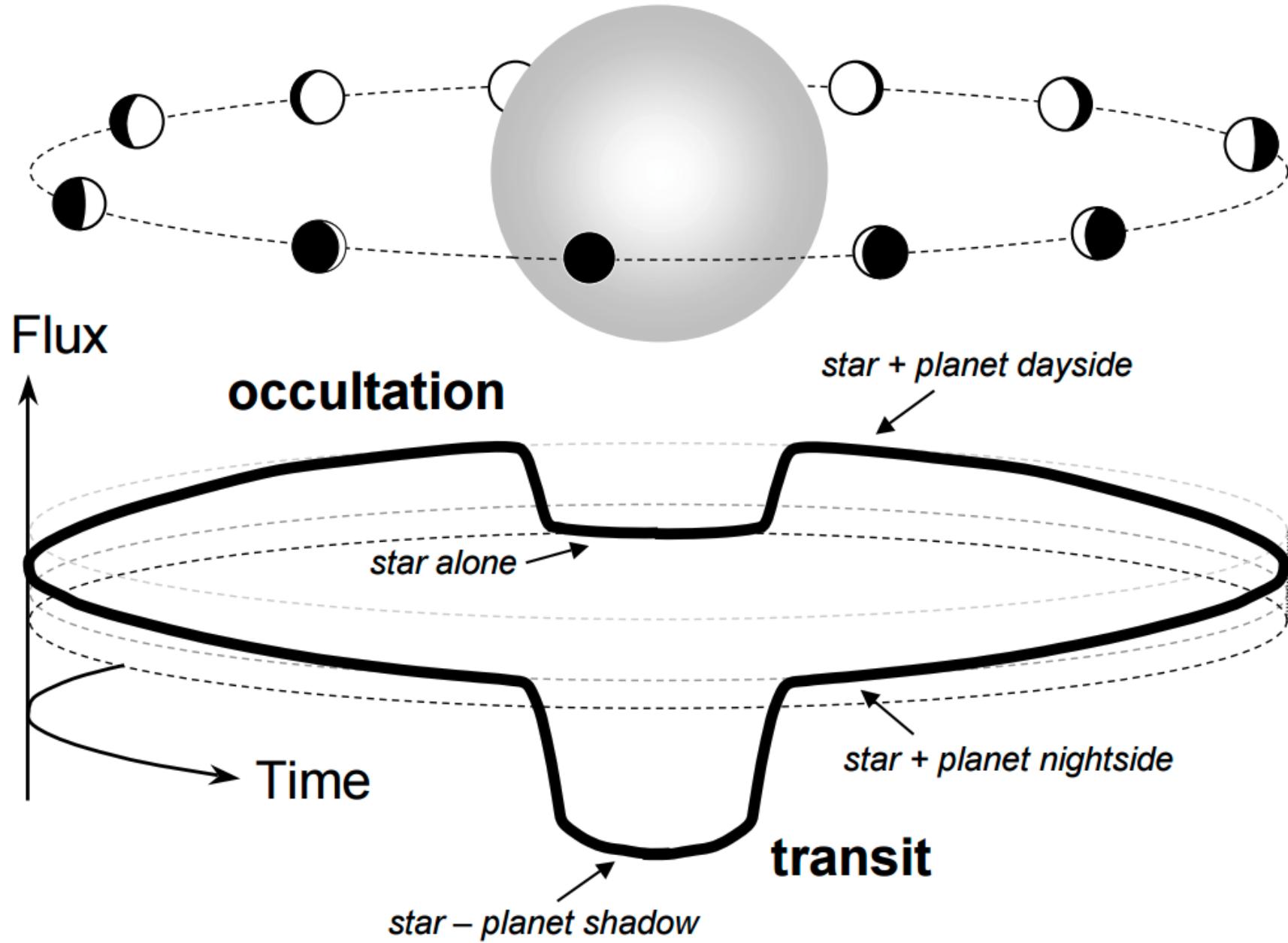


image credit Josh Winn

Earth atmosphere

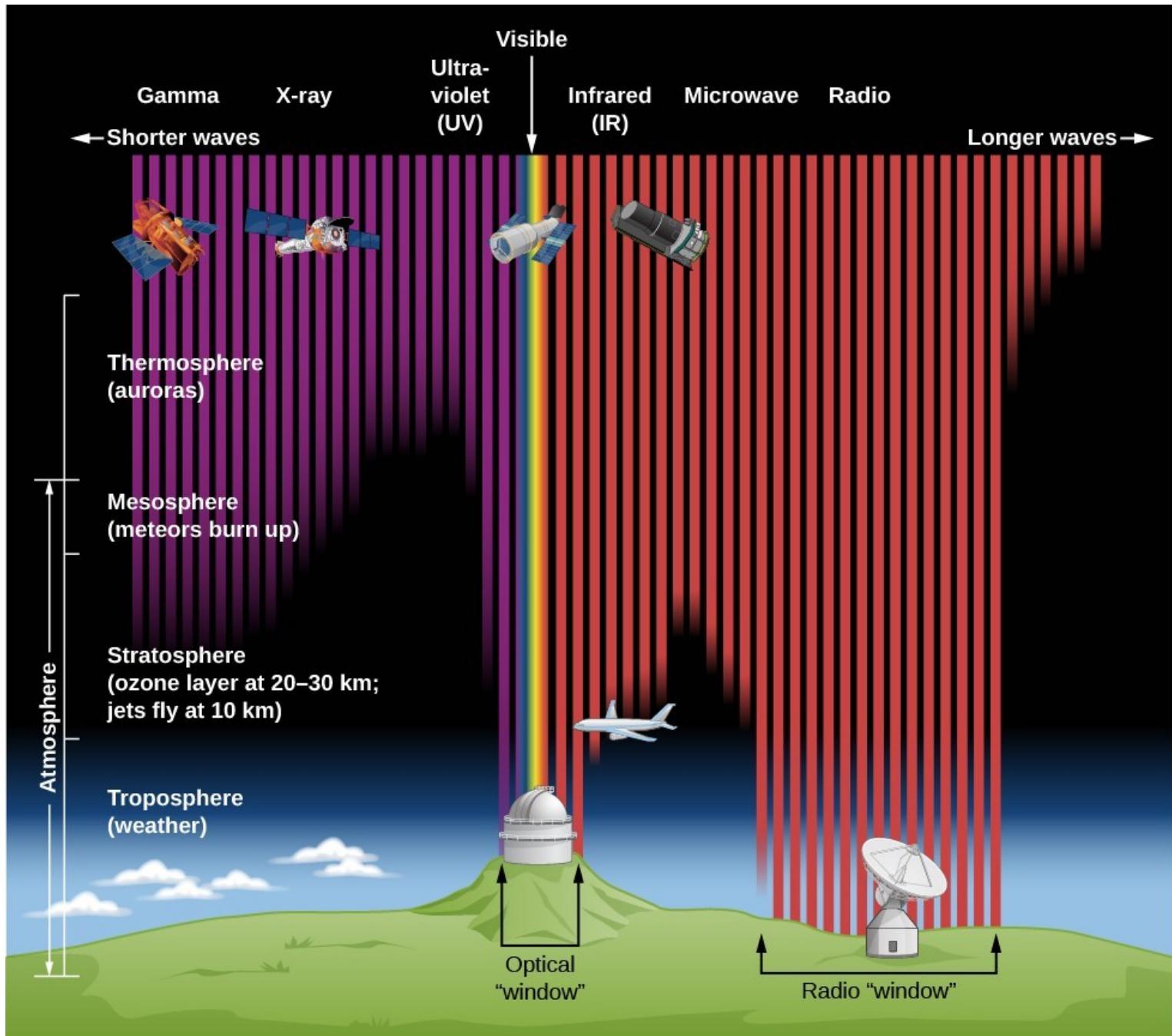


image credit STScI

Atmospheres: spectrophotometry

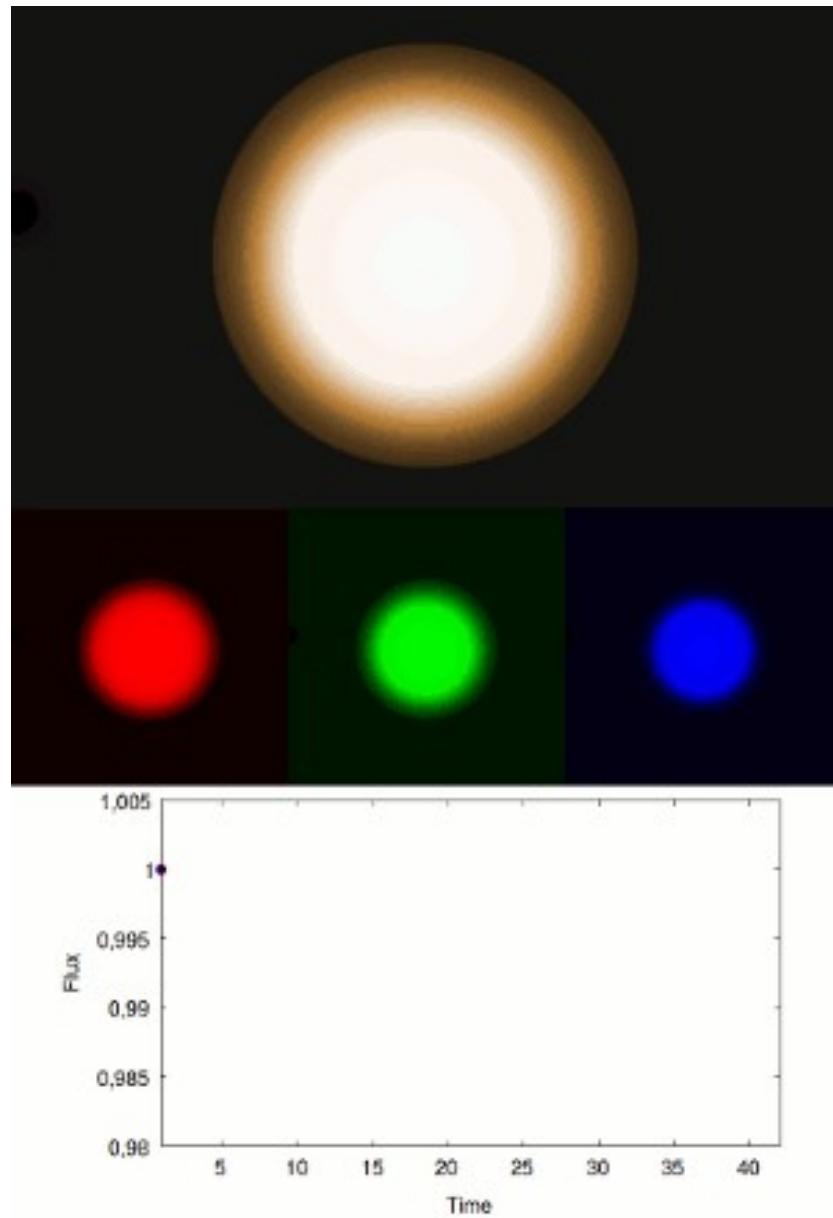
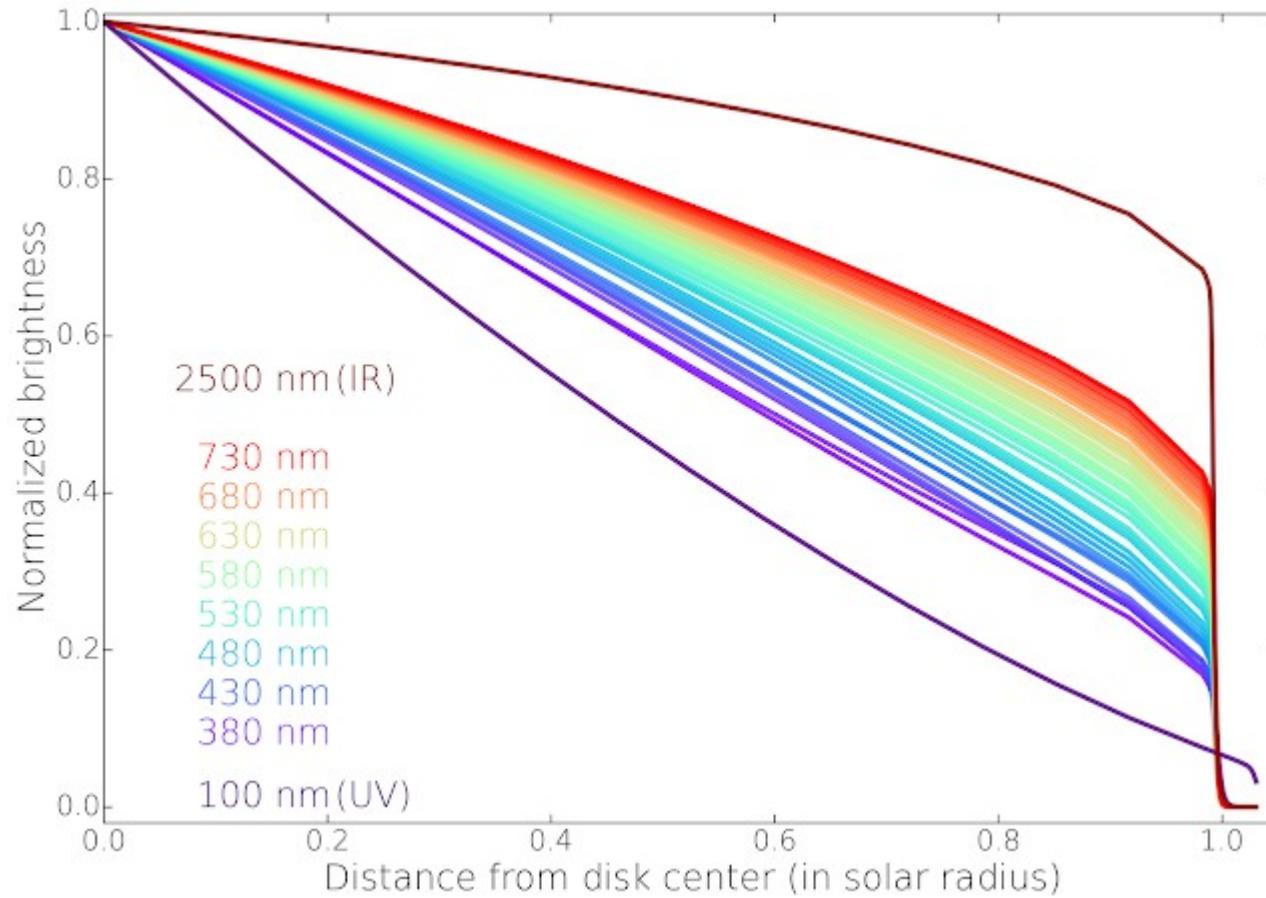


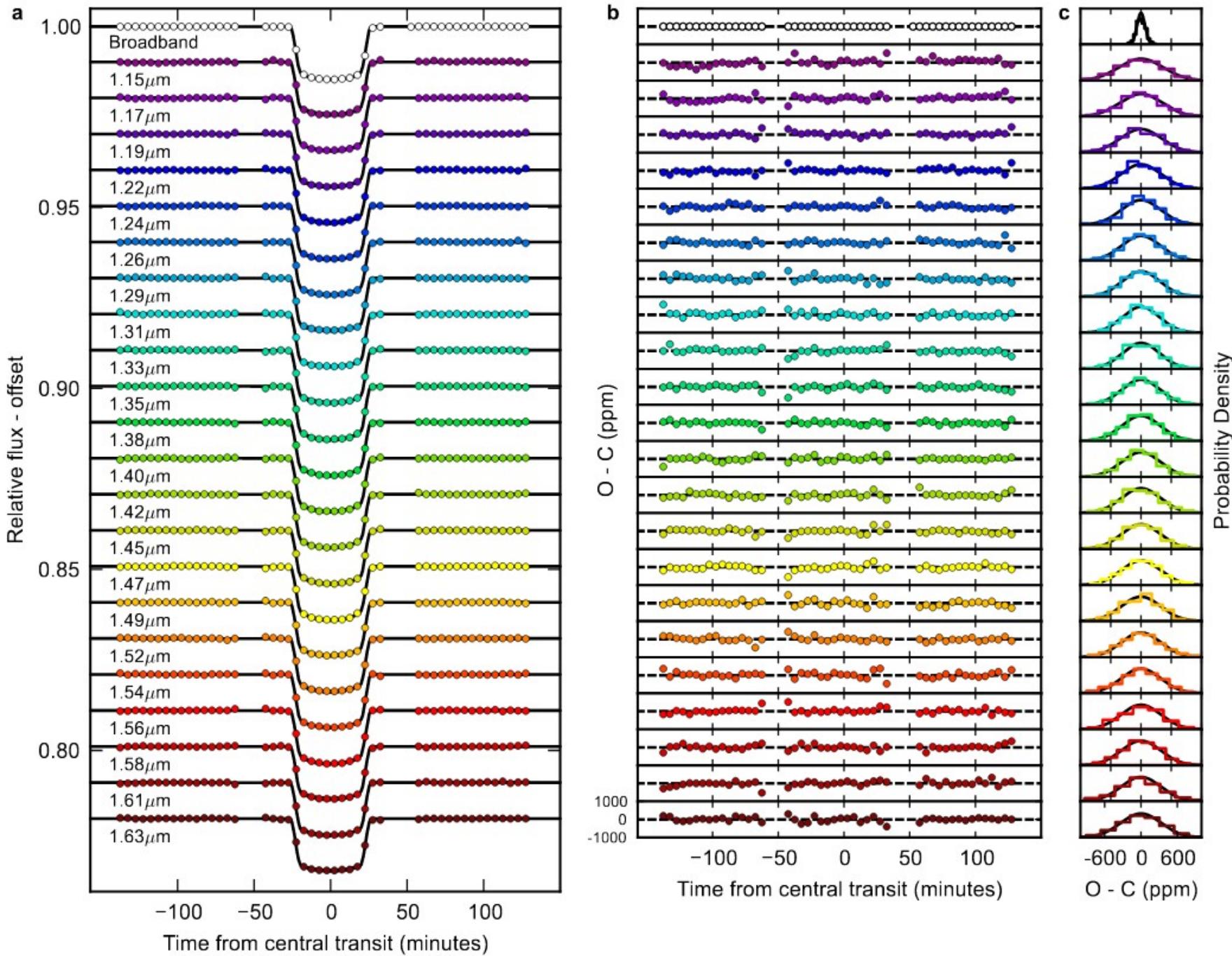
image credit Uppsala University

Atmospheres: spectrophotometry



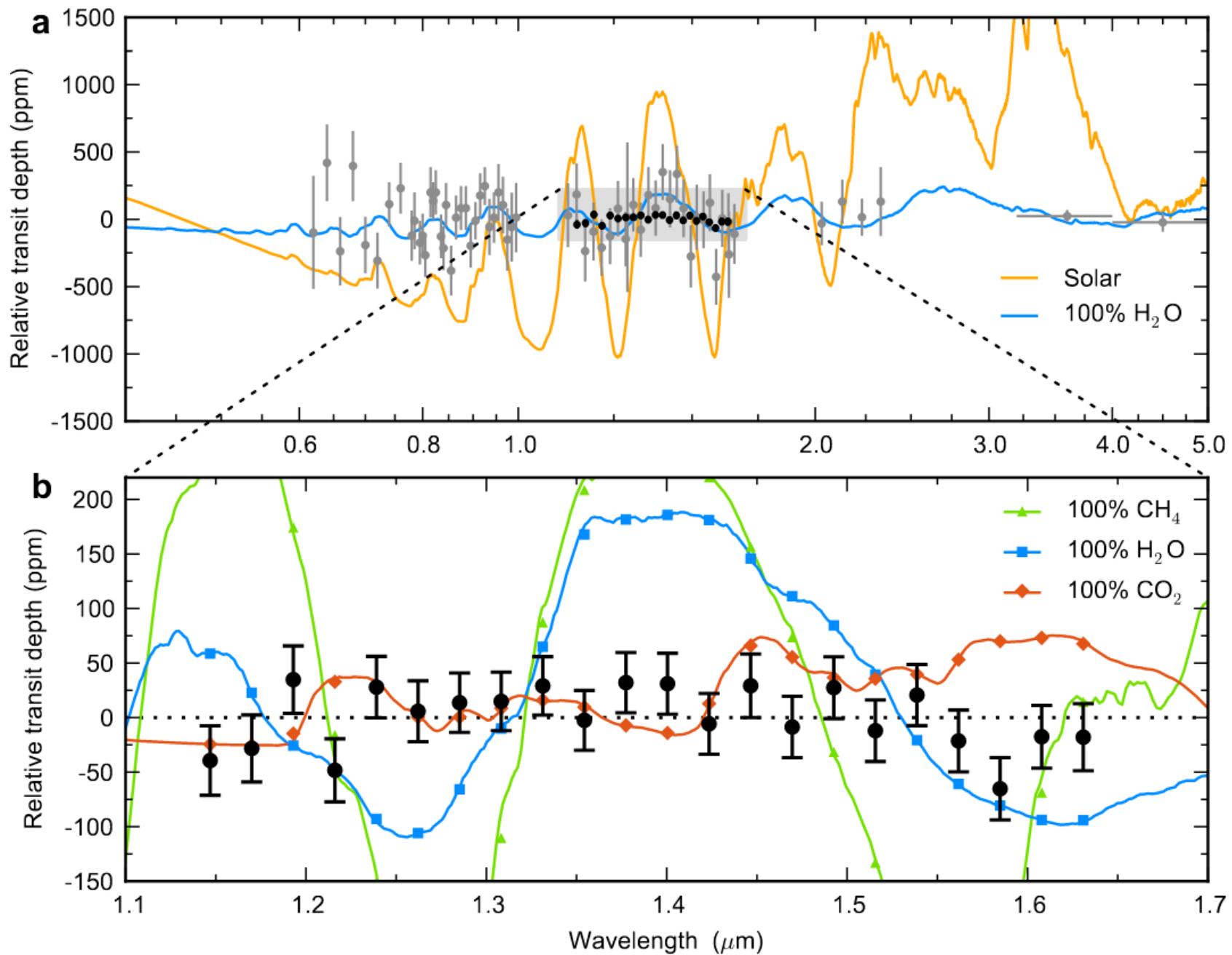
Know your star: limb darkening!

Super-earth GJ 1214



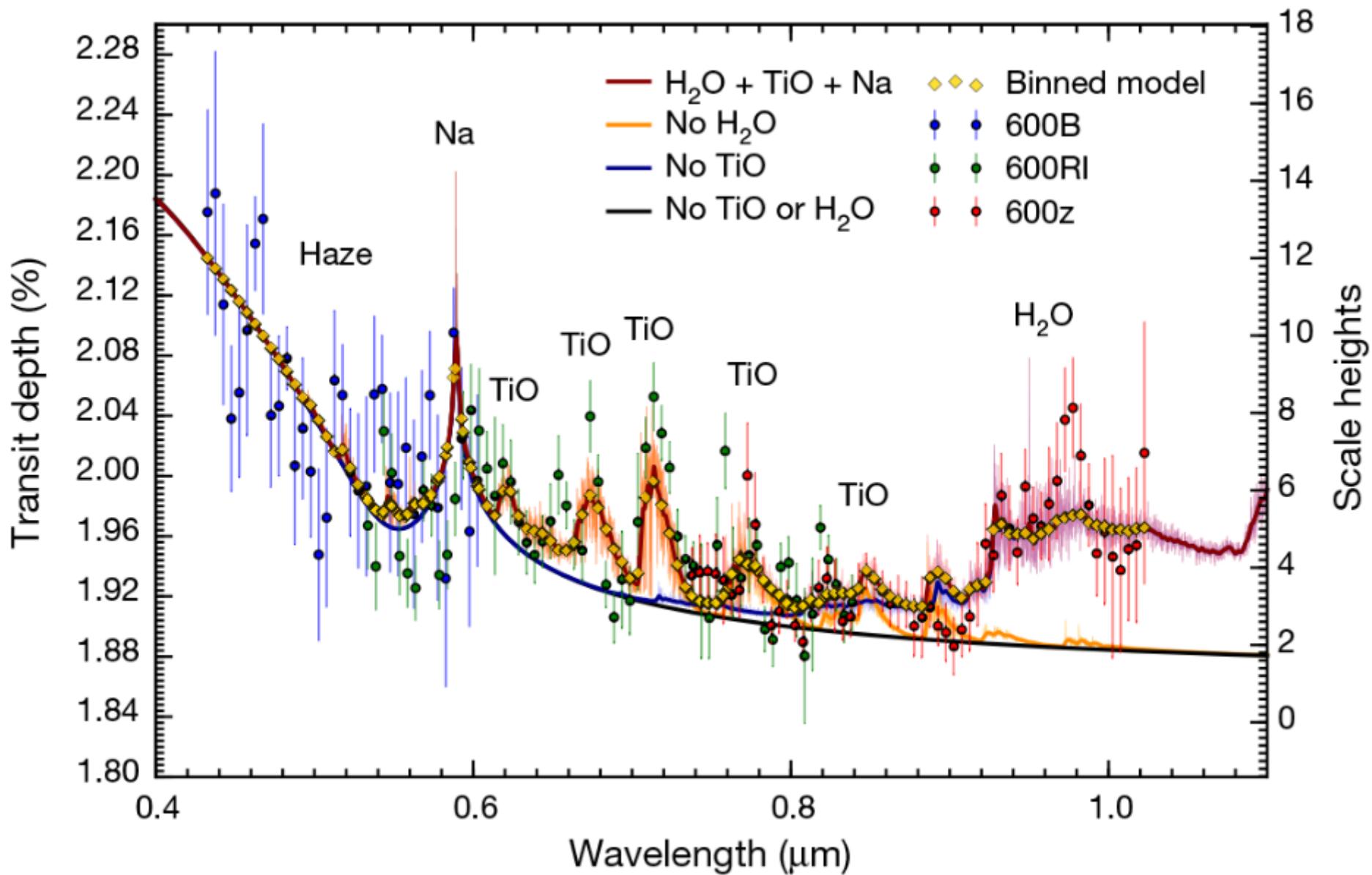
Kreidberg et al. 2014

Super-earth GJ 1214

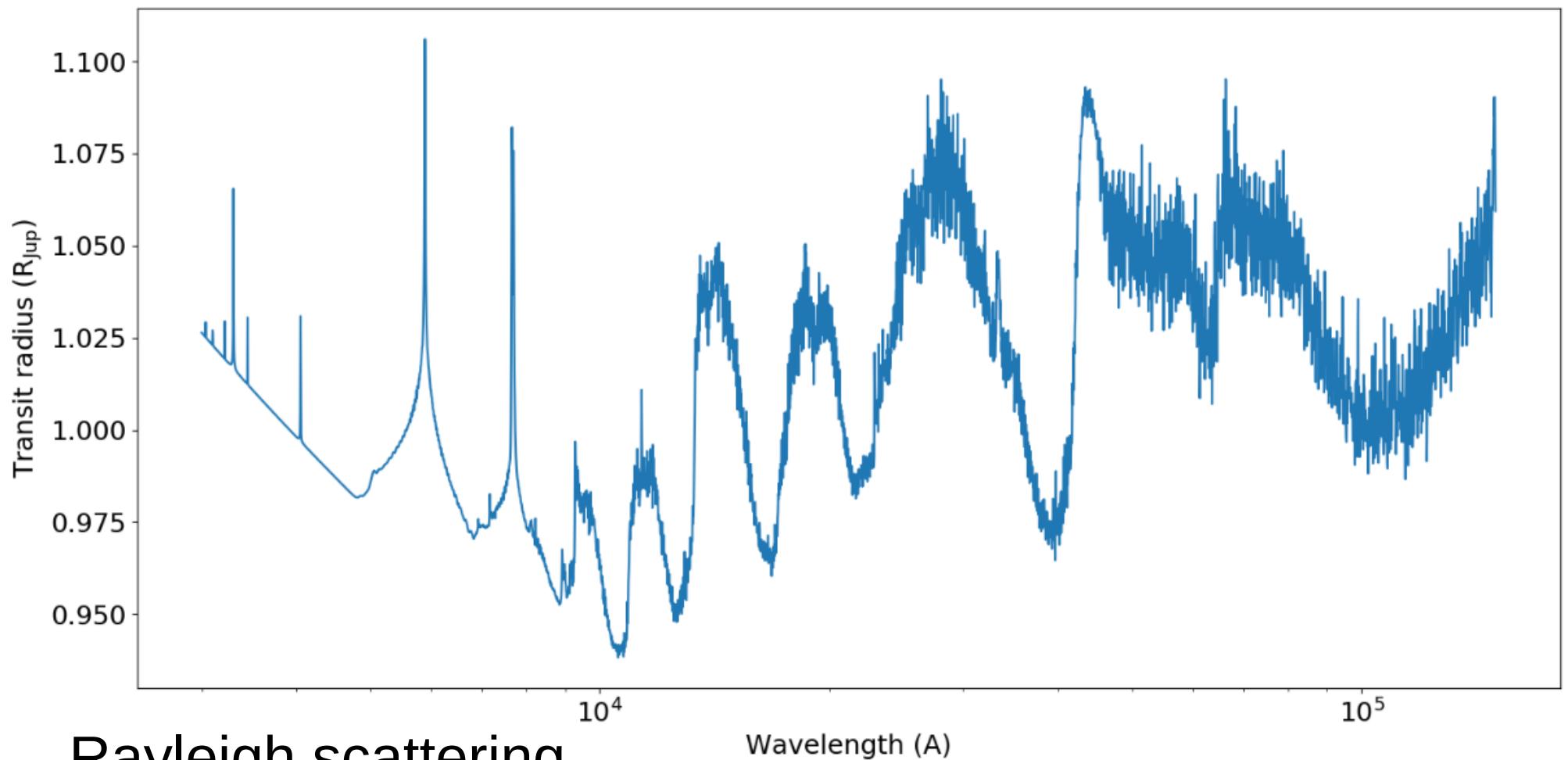


Kreidberg et al. 2014

Hot Jupiter WASP-19b



Hot Jupiter theoretical model (petitRADTRANS)



Rayleigh scattering

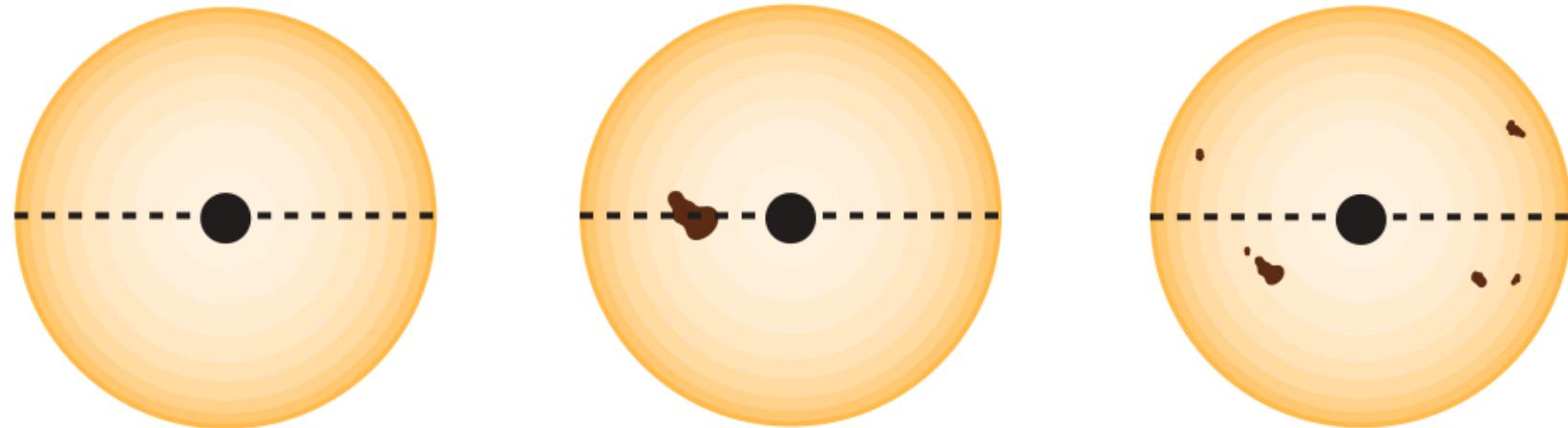
Strong features of alkali metals (Na, K_a)

Molecular bands (H₂O, CO, CO₂, CH₄)

image credit: S. Czesla

Caveat: starspots

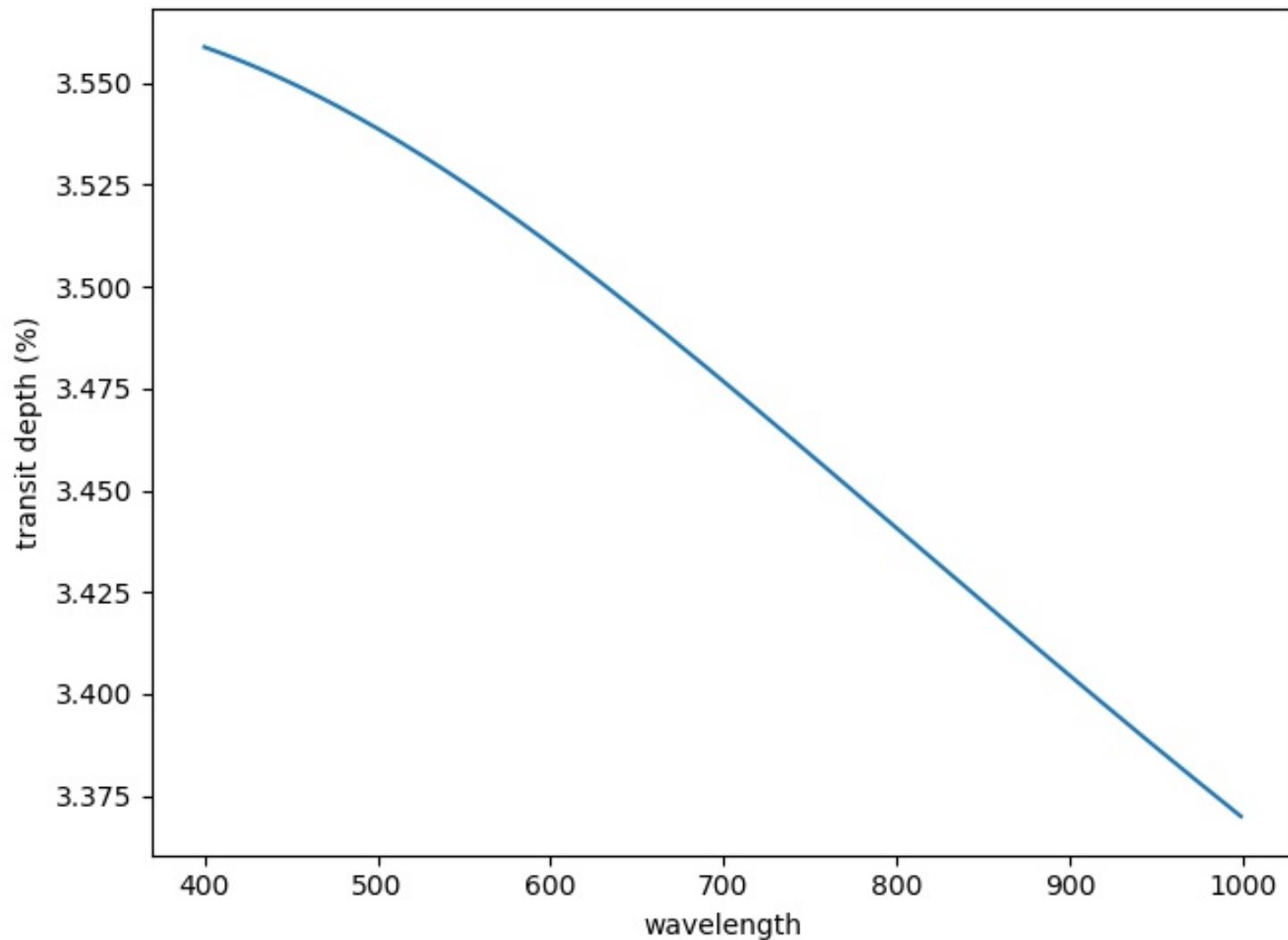
$$\frac{R_p}{R_*} = \sqrt{\delta(1 - \epsilon)}$$



Spot crossing effect and Out Of Transit effect

image credit Helen Giles

Caveat: starspots



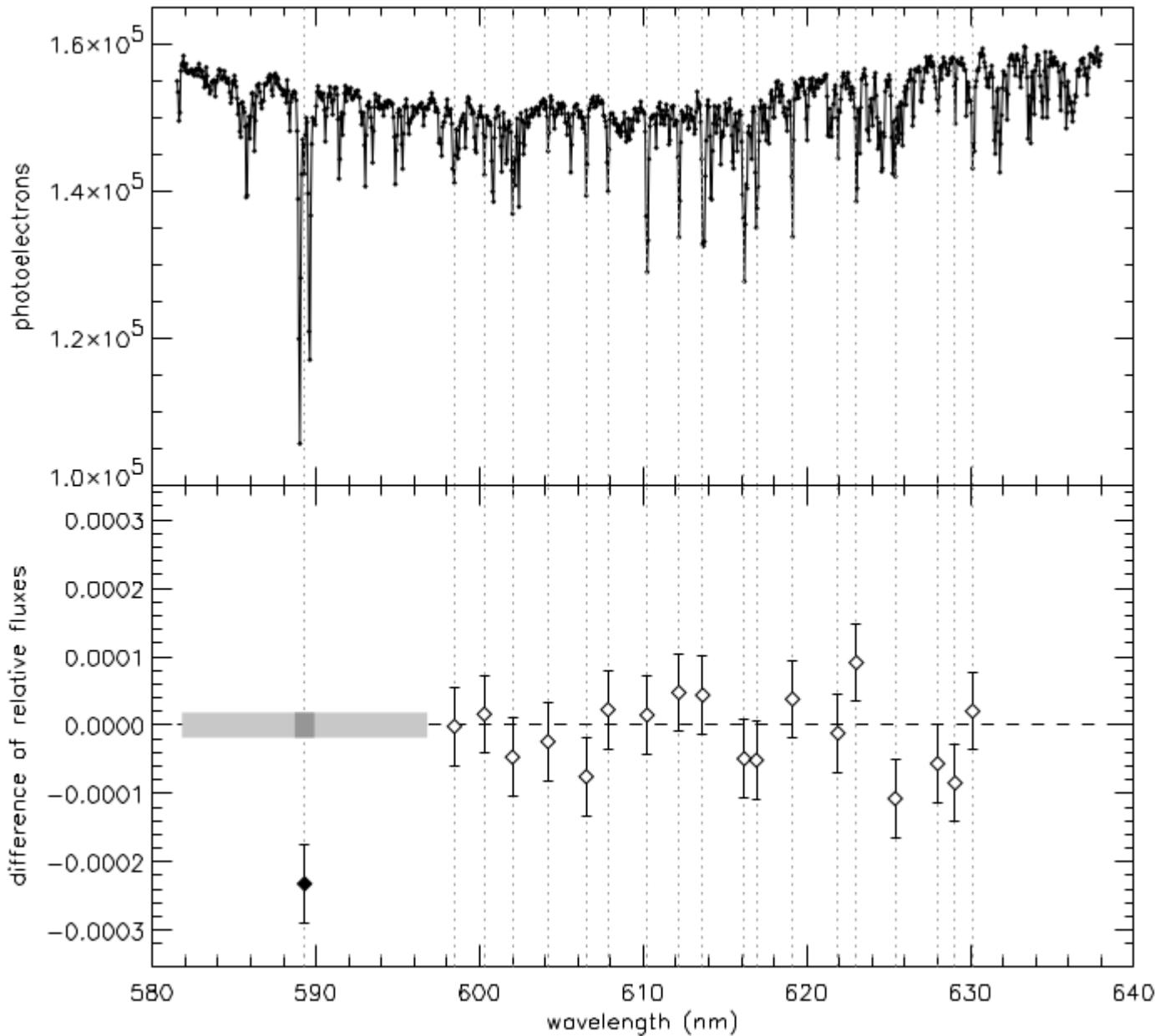
Atmospheres: transmission spectroscopy

Calculate from data:

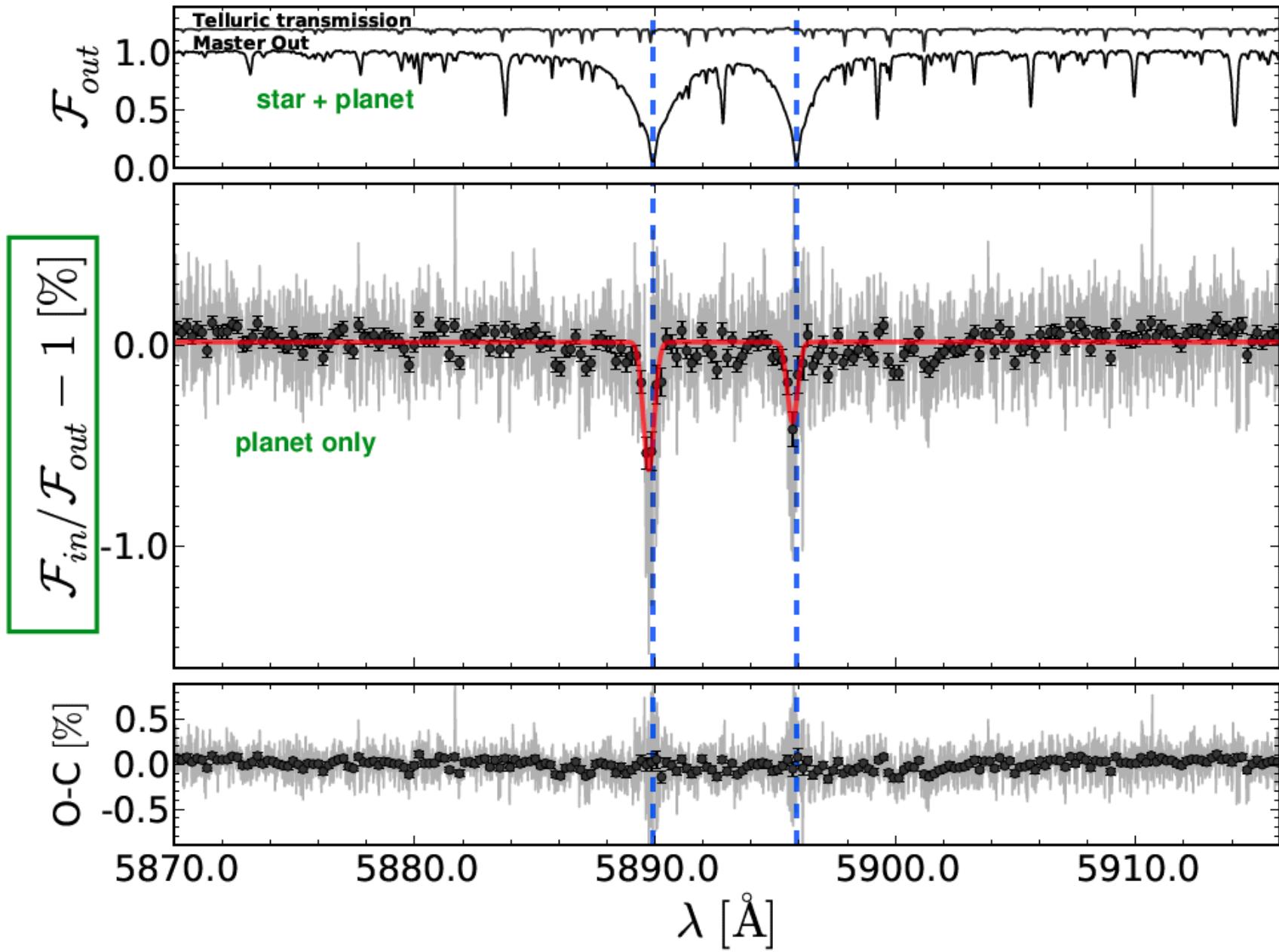
Spectrum during transit / spectrum out of transit

Relative measurement, no absolute flux measurement.
But spectral lines corresponding to absorbers in the
planet's atmosphere will be deeper.

Atmospheres: transmission spectroscopy



Atmospheres: transmission spectroscopy



Atmospheres: transmission spectroscopy

Measures **excess absorption** compared to a spectrally flat (“grey”) transit that reduces the flux by the same fraction everywhere in the spectrum.

